

# DISCOVERY OF A NEW TYPE OF STAR

The background features a dark, textured sky with a gradient from deep purple to a bright orange-red horizon. Below the horizon, a grid of dark lines recedes into the distance, creating a sense of perspective.

# Observations



**The 64 meter Parkes radiotelescope, New South Wales, Australia.**

January 1998 - February 2002



Australia Telescope National Facility  
University of Manchester's Jodrell  
Bank Observatory



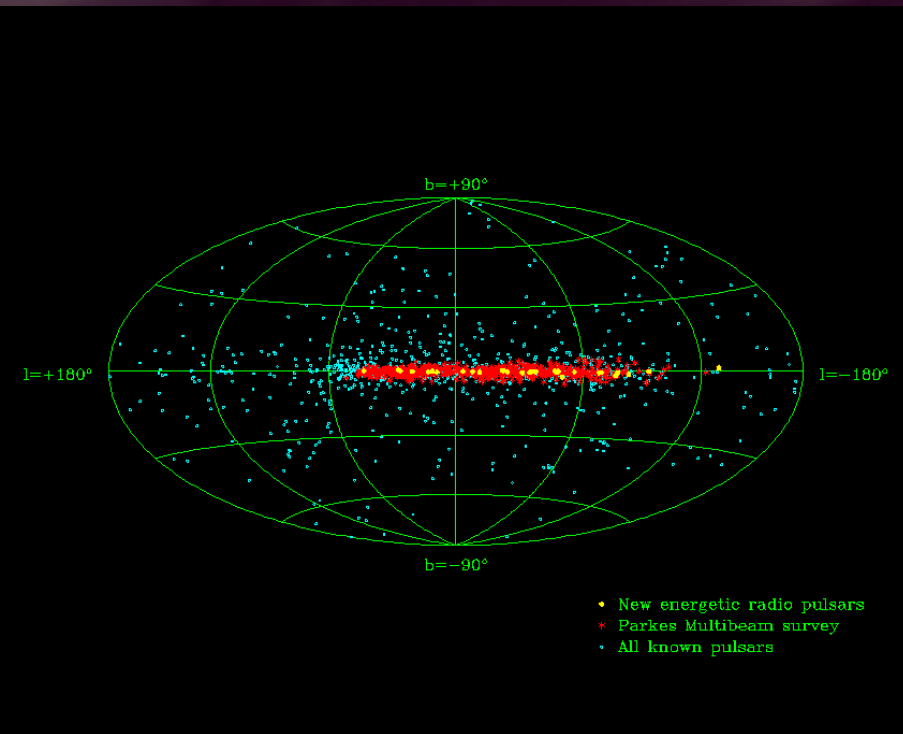
# Observations

- 800 pulsars,
- 11 Rotating Radio Transients (RRATs)

Since August 2003 all the sources (11) have been reobserved.

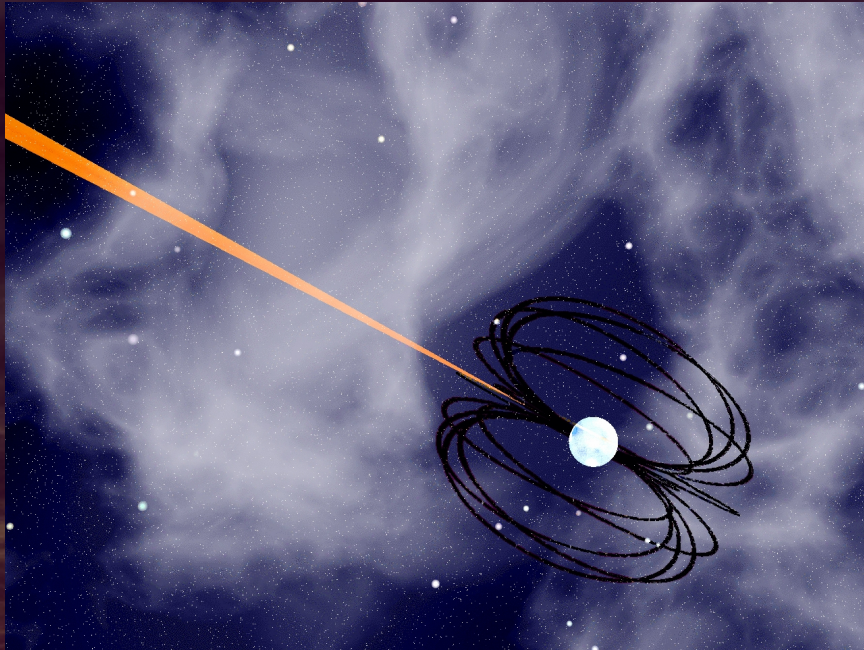
All have shown multiple burst, with between 4 and 229 events detected in total from each object

The density of sources on the sky appears to be greater towards the Galactic plane (8 of 11 have  $|b| < 2^\circ$ )



**McLaughlin, M. A. et al. 2006, Nature**

# Rotating Radio Transients

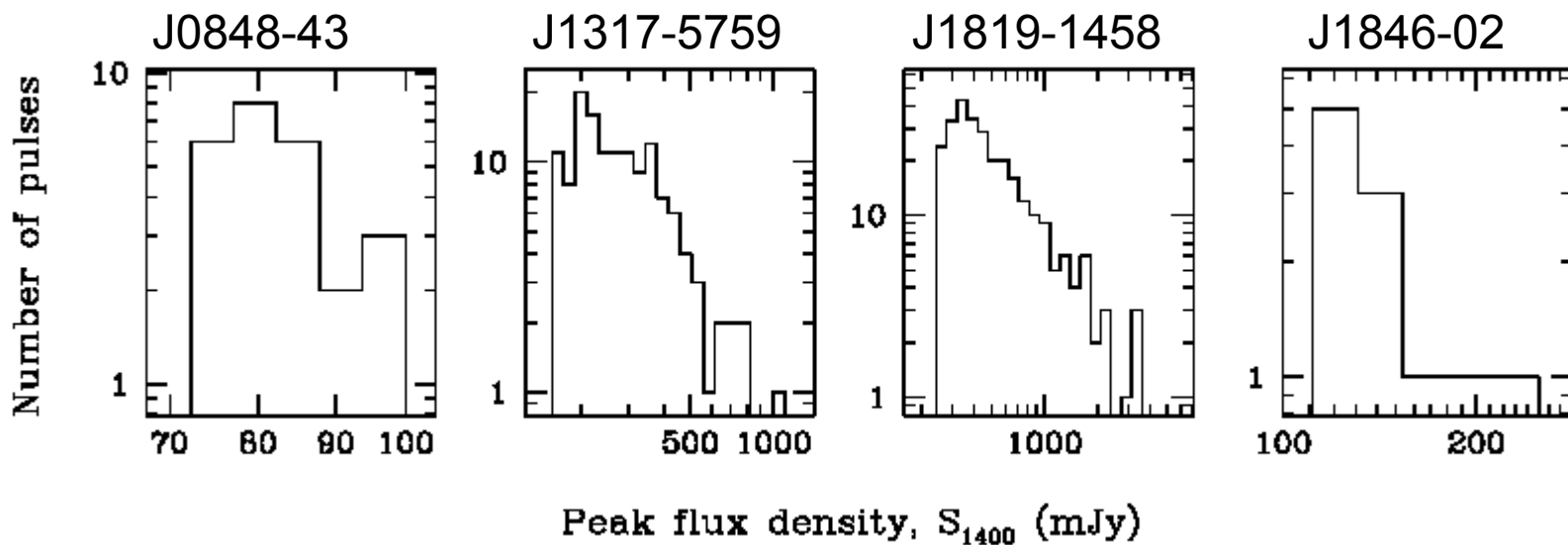
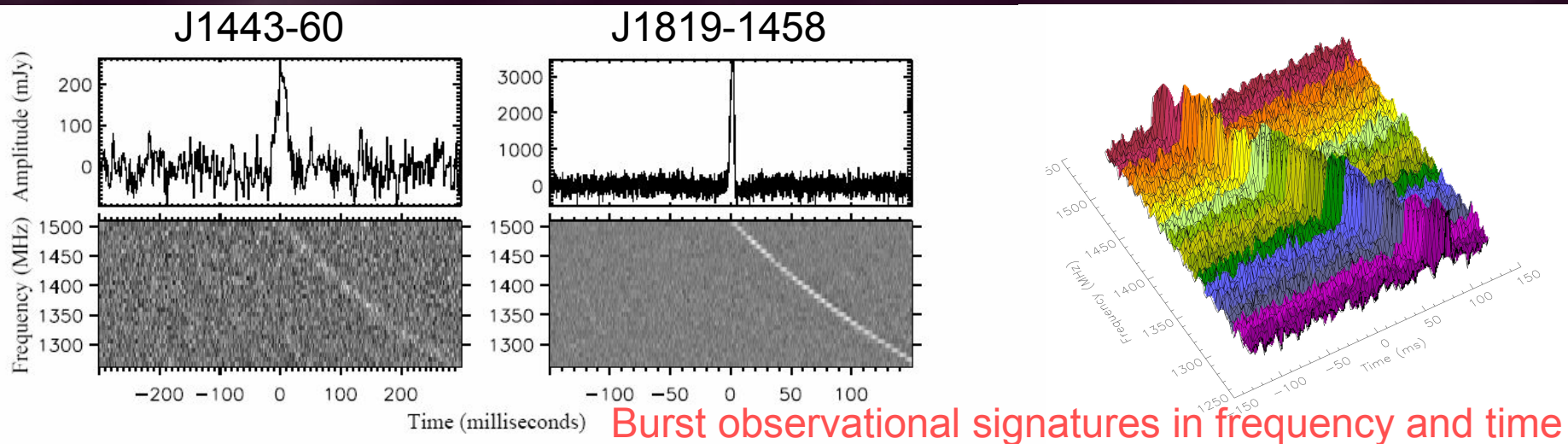


Visualisation of a neutron star, showing the magnetic field lines and the radio beam emanating from a magnetic pole

Durations:	2 – 30 ms
Time interval between burst:	4 min – 3 h
Periodicities:	0.4 – 7 s (for 10 sources)

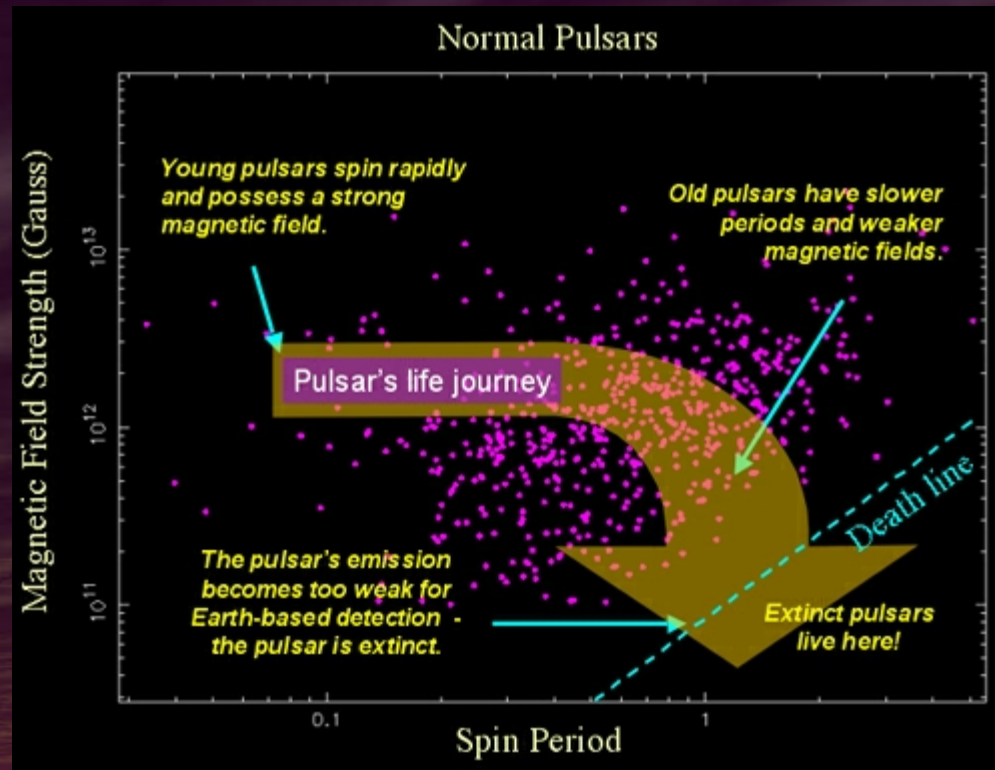
5 of 10  $P > 4$  s,

# Rotating Radio Transients



# Physical interpretation of Rotating Radio Transients

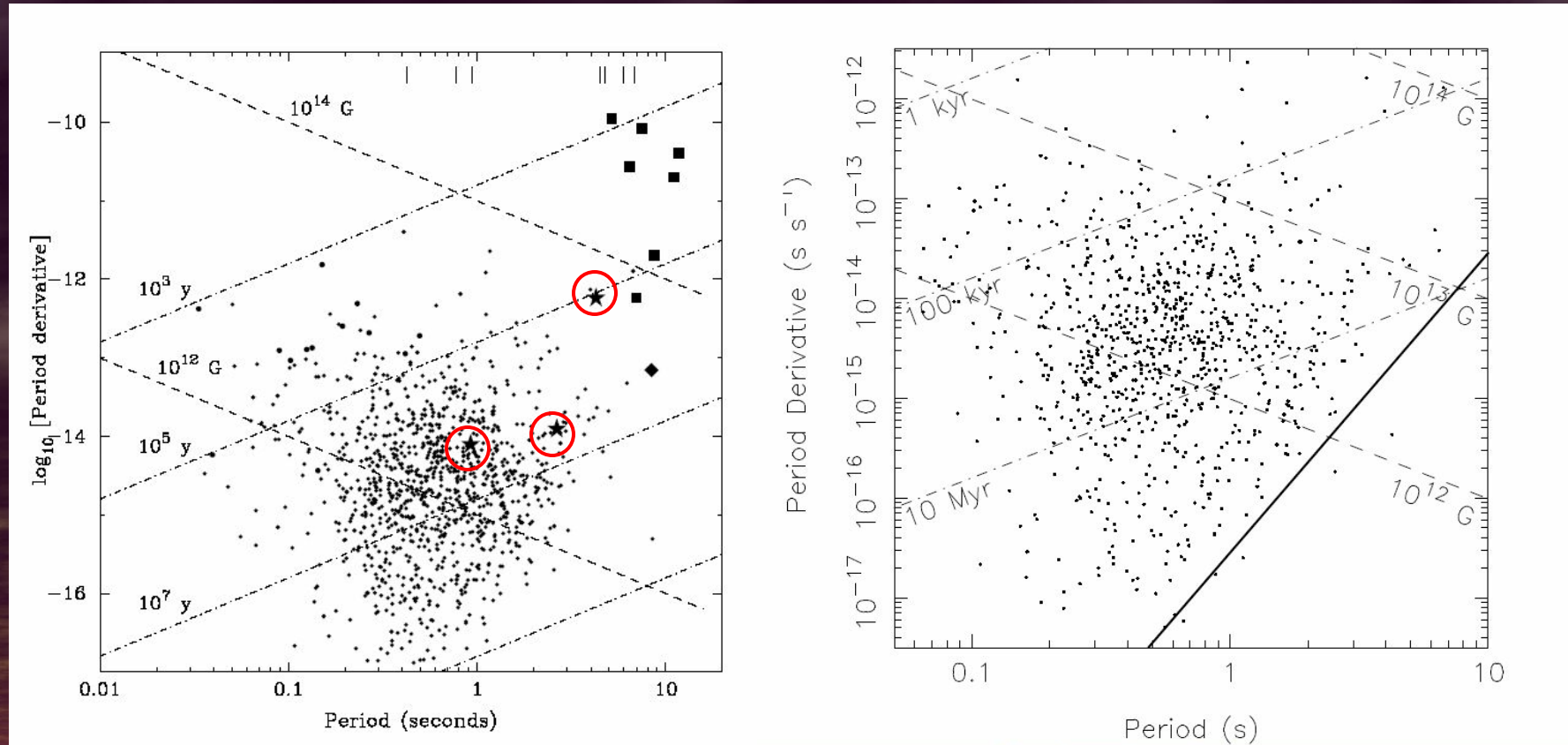
## MODEL I – RE-ACTIVATED DEAD PULSARS



Zhang, B., Gil, J., Dyks, J. 2006, submitted to the Astrophysical Journal Letters

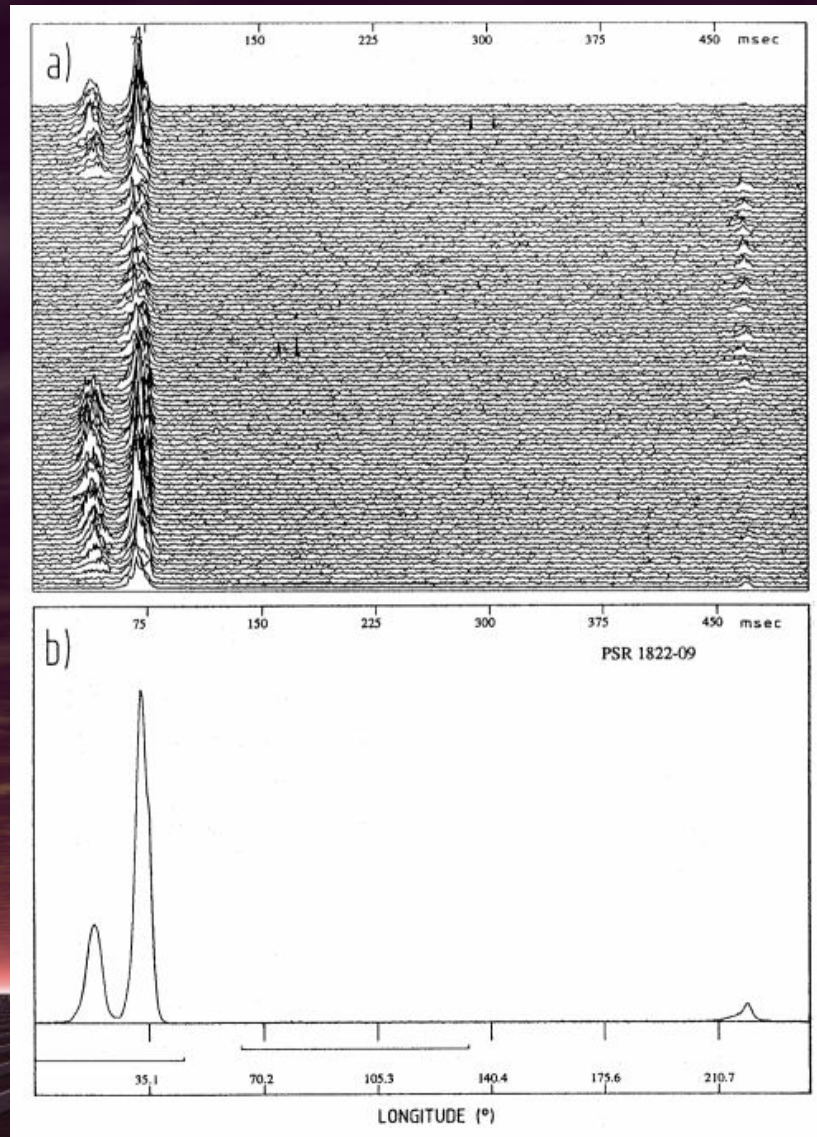
# Physical interpretation of Rotating Radio Transients

## MODEL I – RE-ACTIVATED DEAD PULSARS



# Physical interpretation of Rotating Radio Transients

## MODEL II: REVERSED NORMAL NULLING PULSARS



PSR B1822-09

Gil, J. et al. 1994, A&A,  
282, 45



# Physical interpretation of Rotating Radio Transients

**MODEL : PSR B1822-09**

**Dyks, J. et al. 2005, ApJ, 626, 45**

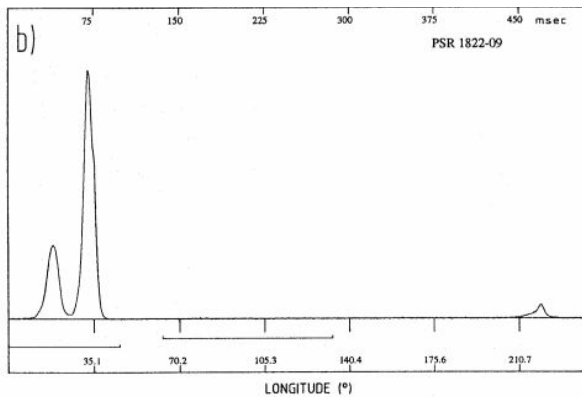
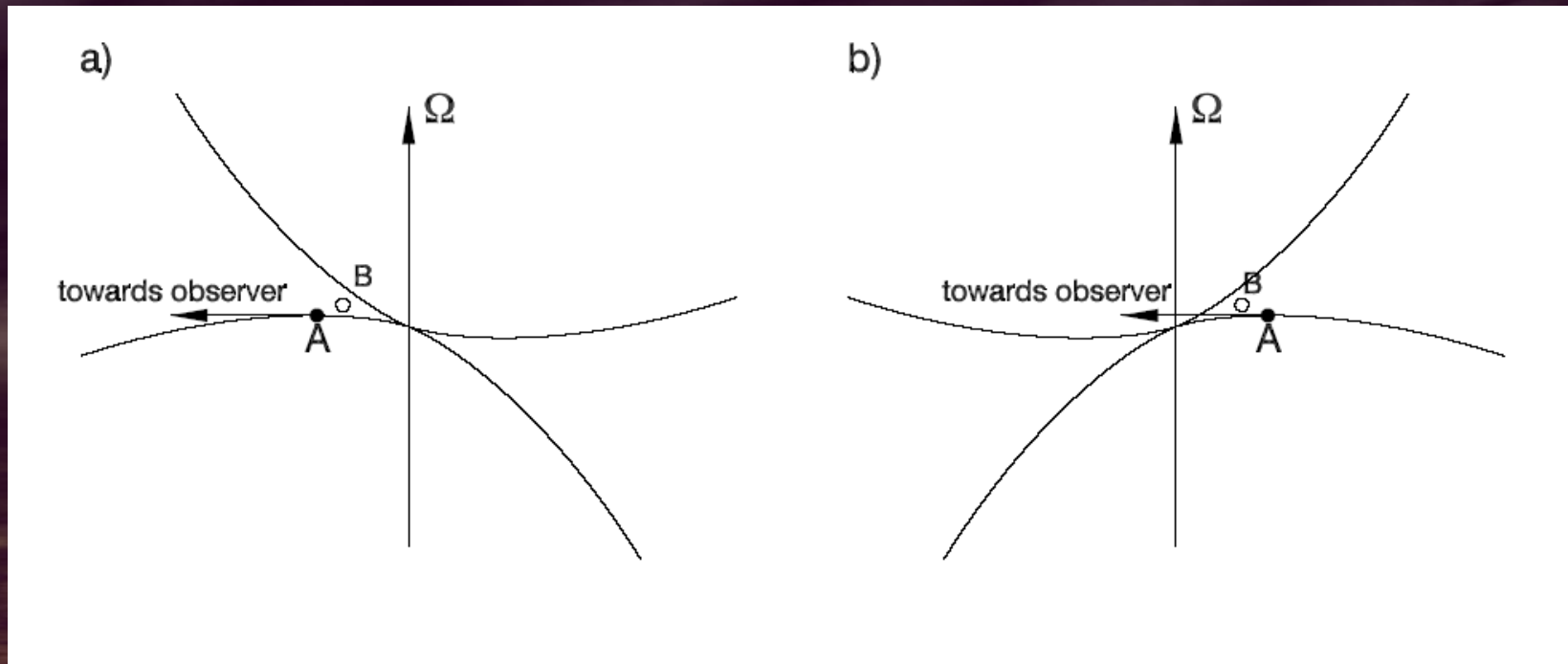
- 1) The mechanism of coherent radio emission must allow radiation into two, opposite, direction.
- 3) The radio waves must be able to propagate through inner regions of the neutron star magnetosphere with strong magnetic field.

The models implies **inward radio emission** in pulsar magnetosphere.

# Physical interpretation of Rotating Radio Transients

## MODEL II: REVERSED NORMAL NULLING PULSARS

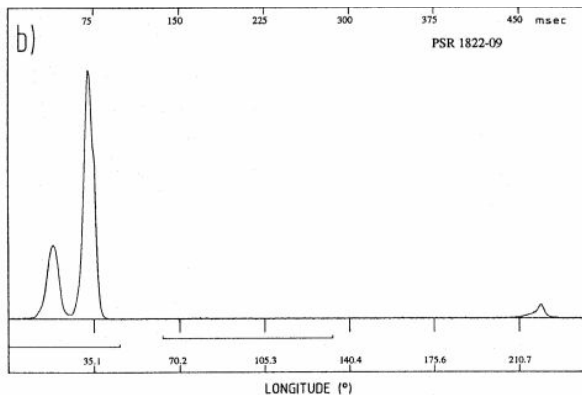
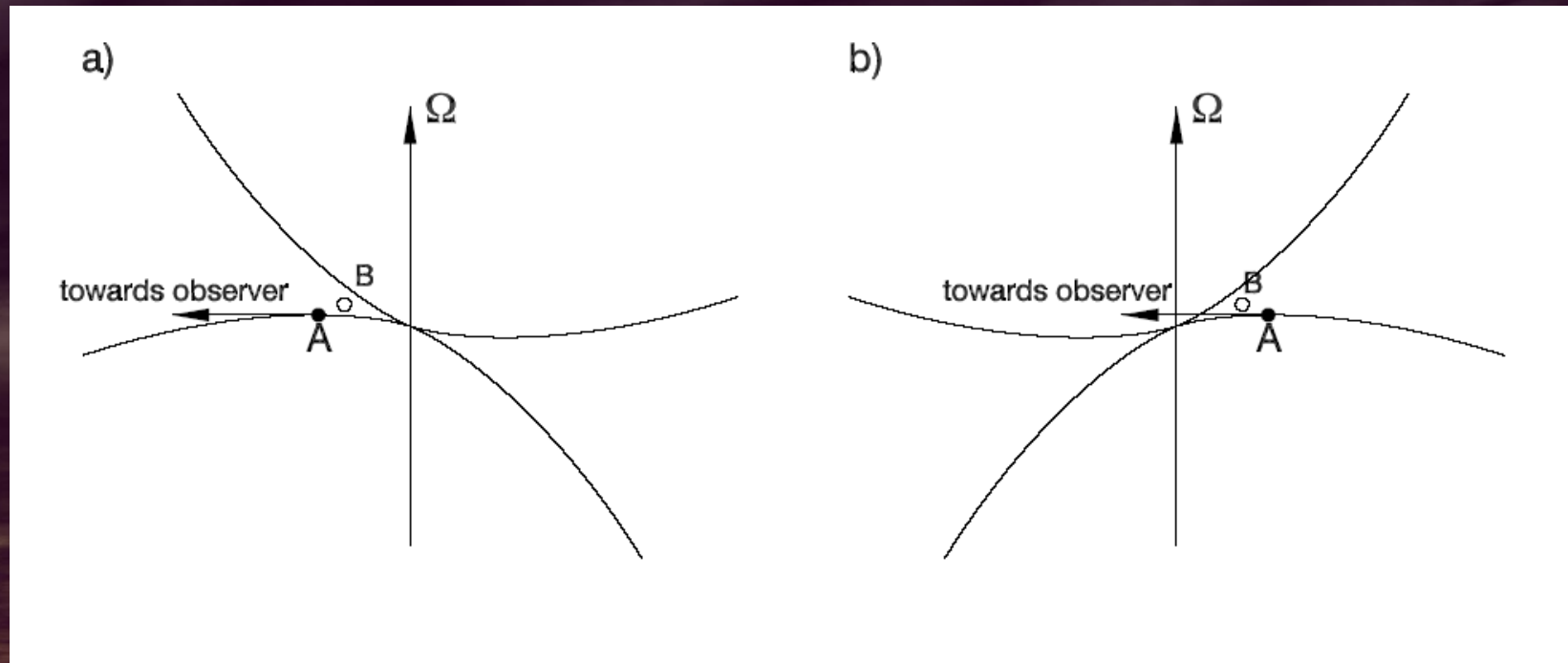
Dyks, J. et al. 2005, ApJ, 626, 45



# Physical interpretation of Rotating Radio Transients

## MODEL II: REVERSED NORMAL NULLING PULSARS

Dyks, J. et al. 2005, ApJ, 626, 45



RRAT – observation of reversed inward-directed pulse component

# X-ray data

## **MODEL I (RE-ACTIVATED DEAD PULSARS)**

**– Strong X-ray emission is not expected, except the thermal component from the cooling neutron star**

## **MODEL II (REVERSED NORMAL NULLING PULSARS)**

**– The X-ray emission properties (luminosity and spectrum) of RRATs should be similar to those of normal, middle-age/old pulsars**

# Observations

**RRAT J1819-1458 was detected with *Chandra* by Reynolds et al. (2006)**

**Observations:  
23, 25, 28 May 2005.**

**Spectrum :**

- **similar to those of comparably-aged radio pulsar,**
- **dominated by a soft thermal component from the cooling neutron star**

**The data are consistent with both models !**